

Statistical Physics

Test 2

mgr Grzegorz Dziewisz

03/06/2026

1. Problem sheet 4. Solve one of the following questions:

- (a) In the presence of an external potential $U(\vec{r})$, the distribution function is inhomogeneous in space:

$$f(\vec{p}, \vec{r}) \propto \exp\left(-\beta\left(\frac{\vec{p}^2}{2m} + U(\vec{r})\right)\right)$$

Consider a gas in the harmonic potential $U(\vec{r}) = \frac{1}{2}\kappa r^2$. Calculate its density $n(\vec{r})$ (up to multiplicative constant).

Notice that mean radius is zero $\langle \vec{r} \rangle = 0$, why? Calculate mean-square radius $\langle r^2 \rangle$.

- (b) The Maxwell-Boltzmann distribution is given as:

$$f(\vec{p}) \propto e^{-\beta\frac{\vec{p}^2}{2m}}$$

Find the distribution function in terms of energy $f_E(E)$ and express the n^{th} energy moment of this distribution $\langle E^n \rangle$, $n \in \mathbb{N}$.

Hint:

$$\Gamma(x) = \int_0^\infty t^{x-1} e^{-t} dt$$

2. Problem sheet 5 - partition function. Solve one of the following questions:

- (a) The one-dimensional quantum harmonic oscillator has its energy spectrum given by:

$$E_n = \left(n + \frac{1}{2}\right) \hbar\omega$$

where $\omega > 0$ is the angular frequency of the oscillator, \hbar is Dirac constant, $n \in \mathbb{N}_0$. Find the partition function for such an oscillator and then calculate its internal energy and heat capacity.

Hint:

$$Z = \frac{e^{-\frac{1}{2}\beta\hbar\omega}}{1 - e^{-\beta\hbar\omega}}$$

- (b) Consider a system composed of two subsystems A , B . In general, its energy states can be written as:

$$E_{ij} = E_i^A + E_j^B + V_{ij}$$

where V_{ij} corresponds to interactions - if it vanishes, these subsystems are independent. It is reasonable to conclude then that other properties like **free energy** or **entropy** also can be decomposed into a sum over these subsystems. In statistical mechanics, your task is to prove this.

Hint: start with

$$Z_{AB} = \sum_{i,j} e^{-\beta E_{ij}}$$

Examples of solutions

1. Problem sheet 4

(a) Of course, since the potential does not depend on the momenta of particles:

$$\begin{aligned} n(\vec{r}) &= \int_{\mathbb{R}^3} f(\vec{p}, \vec{r}) d^3p \\ &= C \exp(-\beta U(\vec{r})) \int_{\mathbb{R}^3} \exp\left(-\frac{\beta \vec{p}^2}{2m}\right) d^3p \\ &= \tilde{C} \exp\left(-\frac{1}{2}\beta\kappa \vec{r}^2\right) \end{aligned}$$

Since the integral is a Gaussian one, hence it's finite. One can easily notice that:

$$\tilde{C} = n(\vec{0}) \equiv n_0$$

Mean radius of this density function is zero since, by construction, the potential is centred at the beginning of the \mathbb{R}^3 . In other words:

$$\begin{aligned} \langle x \rangle &= n_0 \int_{\mathbb{R}^3} x e^{-\frac{1}{2}\beta\kappa(x^2+y^2+z^2)} dx dy dz \\ &= \int_{\mathbb{R}} e^{-\frac{1}{2}\beta\kappa y^2} dy \cdot \int_{\mathbb{R}} e^{-\frac{1}{2}\beta\kappa z^2} dz \cdot \int_{\mathbb{R}} x e^{-\frac{1}{2}\beta\kappa x^2} dx \end{aligned}$$

and since x is an odd function, the Gaussian is even, the total integral vanishes. The same applies for other components of $\vec{r} = (x, y, z)$

In order to calculate mean-square radius:

$$\begin{aligned} \langle r^2 \rangle &= \langle x^2 + y^2 + z^2 \rangle \\ &= n_0 \int_{\mathbb{R}^3} (x^2 + y^2 + z^2) e^{-\frac{1}{2}\beta\kappa(x^2+y^2+z^2)} dx dy dz \\ &= n_0 \int_{\mathbb{R}} x^2 e^{-\frac{1}{2}\beta\kappa x^2} dx \cdot \int_{\mathbb{R}} e^{-\frac{1}{2}\beta\kappa y^2} dy \cdot \int_{\mathbb{R}} e^{-\frac{1}{2}\beta\kappa z^2} dz \cdot 3 \end{aligned}$$

where the factor 3 follows from the separation of the original integral into a sum of products of integrals and the indistinguishability of integrals.

It is useful to remember that:

$$I(\alpha) = \int_{\mathbb{R}} e^{-\alpha x^2} dx = \sqrt{\frac{\pi}{\alpha}}, \quad \alpha > 0$$

This gives us two integrals from above. We can use this also to calculate remaining one:

$$\begin{aligned} \int_{\mathbb{R}} x^2 e^{-\alpha x^2} dx &= - \int_{\mathbb{R}} \frac{\partial}{\partial \alpha} e^{-\alpha x^2} dx = - \frac{d}{d\alpha} \int_{\mathbb{R}} e^{-\alpha x^2} dx \\ &= - \frac{d}{d\alpha} I(\alpha) = -\sqrt{\pi} \frac{d}{d\alpha} \left(\alpha^{-\frac{1}{2}}\right) \\ &= -\sqrt{\pi} \cdot \left(-\frac{1}{2}\right) \alpha^{-\frac{3}{2}} = \frac{\sqrt{\pi}}{2\alpha^{3/2}} \end{aligned}$$

For our purpose: $\alpha = \frac{1}{2}\beta\kappa$. As a result:

$$\begin{aligned} \langle r^2 \rangle &= \dots = 3n_0 \cdot \frac{\sqrt{\pi}}{2\left(\frac{1}{2}\beta\kappa\right)^{3/2}} \cdot \sqrt{\frac{\pi}{\frac{1}{2}\beta\kappa}} \cdot \sqrt{\frac{\pi}{\frac{1}{2}\beta\kappa}} \\ &= 6\sqrt{2} n_0 \pi^{3/2} \kappa^{-5/2} \beta^{-5/2} \end{aligned}$$

(b) Since f is a distribution, then:

$$\int f(\vec{p}) d^3p = \int f_E(E) dE$$

The energy (kinetic energy) of a gas is:

$$E(\vec{p}) = \frac{\vec{p}^2}{2m} = \frac{p^2}{2m}$$

where $p = |\vec{p}|$ is the length of momentum vector. From this:

$$p^2 = 2mE, \quad 2pdp = 2mdE \implies dp = \frac{m}{p(E)} dE = \frac{m}{\sqrt{2mE}} dE = \sqrt{\frac{m}{2E}} dE$$

Let's put everything together:

$$\begin{aligned} \int f_E(E) dE &= \int f(\vec{p}) d^3p \\ &= C \int e^{-\beta E(\vec{p})} d^3p \\ &= C \int d\phi \int \sin(\theta) d\theta \int e^{-\beta E(p)} p^2 dp \\ &= \tilde{C} \int e^{-\beta E} \cdot 2mE \cdot \sqrt{\frac{m}{2E}} dE \\ &= \bar{C} \int \sqrt{E} e^{-\beta E} dE \end{aligned}$$

and as a result:

$$f_E(E) \propto \sqrt{E} e^{-\beta E}$$

and \bar{C} is just normalisation constant.

To calculate the n^{th} moment of this distribution:

$$\begin{aligned} \langle E^n \rangle &= \int_0^\infty E^n f_E(E) dE \\ &= \bar{C} \int_0^\infty E^n \sqrt{E} e^{-\beta E} dE \\ &= \bar{C} \int_0^\infty E^{n+\frac{1}{2}} e^{-\beta E} dE \\ &\left| \begin{array}{l} \beta E = t \implies \beta dE = dt \implies dE = \frac{1}{\beta} dt \\ E^{n+\frac{1}{2}} = (\beta E)^{n+\frac{1}{2}} \cdot \frac{1}{\beta^{n+\frac{1}{2}}} = \frac{t^{n+\frac{1}{2}}}{\beta^{n+\frac{1}{2}}} \\ E = 0 \implies t = 0 \\ E \rightarrow \infty \implies t \rightarrow \infty \end{array} \right. \\ &= \bar{C} \int_0^\infty \frac{1}{\beta^{n+\frac{3}{2}}} t^{n+\frac{1}{2}} e^{-t} dt \\ &= \frac{\bar{C}}{\beta^{n+\frac{3}{2}}} \Gamma\left(n + \frac{3}{2}\right) \end{aligned}$$

$$\langle E^n \rangle = \frac{\bar{C}}{\beta^{n+\frac{3}{2}}} \Gamma\left(n + \frac{3}{2}\right)$$

2. Problem sheet 5.

(a) First, let us calculate the partition function:

$$\begin{aligned}
 Z &= \sum_{n \in \mathbb{N}_0} e^{-\beta E_n} = \sum_{n=0}^{\infty} e^{-\beta(n+\frac{1}{2})\hbar\omega} \\
 &= \sum_{n=0}^{\infty} e^{-\beta n\hbar\omega} e^{-\frac{1}{2}\beta\hbar\omega} = e^{-\frac{1}{2}\beta\hbar\omega} \sum_{n=0}^{\infty} (e^{-\beta\hbar\omega})^n \\
 &= e^{-\frac{1}{2}\beta\hbar\omega} \cdot \frac{1}{1 - e^{-\beta\hbar\omega}}
 \end{aligned}$$

where we used the sum for an infinite geometric series with ratio $|q| < 1$.
In order to calculate the internal energy:

$$\begin{aligned}
 \ln(Z) &= \ln\left(e^{-\frac{1}{2}\beta\hbar\omega} \cdot \frac{1}{1 - e^{-\beta\hbar\omega}}\right) \\
 &= \ln\left(e^{-\frac{1}{2}\beta\hbar\omega}\right) - \ln\left(1 - e^{-\beta\hbar\omega}\right) \\
 &= -\frac{1}{2}\beta\hbar\omega - \ln\left(1 - e^{-\beta\hbar\omega}\right)
 \end{aligned}$$

$$\begin{aligned}
 U &= -\frac{\partial \ln(Z)}{\partial \beta} \\
 &= -\frac{\partial}{\partial \beta} \left(-\frac{1}{2}\beta\hbar\omega - \ln\left(1 - e^{-\beta\hbar\omega}\right) \right) \\
 &= \frac{1}{2}\hbar\omega + \frac{\hbar\omega e^{-\beta\hbar\omega}}{1 - e^{-\beta\hbar\omega}}
 \end{aligned}$$

To calculate heat capacity, we know that:

$$\beta = \frac{1}{k_B T} \implies \frac{d\beta}{dT} = -\frac{1}{k_B T^2} = -k_B \beta^2$$

Therefore:

$$\begin{aligned}
 C_v &= \frac{\partial U}{\partial T} = \frac{\partial U}{\partial \beta} \frac{d\beta}{dT} \\
 &= -k_B \beta^2 \frac{\partial U}{\partial \beta} \\
 &= -k_B \beta^2 \frac{\partial}{\partial \beta} \left(\frac{\hbar\omega e^{-\beta\hbar\omega}}{1 - e^{-\beta\hbar\omega}} \right) \\
 &= -k_B \beta^2 \left(\frac{-\hbar^2 \omega^2 e^{-\beta\hbar\omega}}{1 - e^{-\beta\hbar\omega}} + \frac{\hbar\omega e^{-\beta\hbar\omega} \cdot (-1) \cdot (-1) \cdot (-\hbar\omega) e^{-\beta\hbar\omega}}{(1 - e^{-\beta\hbar\omega})^2} \right) \\
 &= k_B \beta^2 \hbar^2 \omega^2 e^{-\beta\hbar\omega} \frac{1}{(1 - e^{-\beta\hbar\omega})^2} (1 - e^{-\beta\hbar\omega} + e^{-\beta\hbar\omega}) \\
 &= k_B \beta^2 \hbar^2 \omega^2 \frac{e^{-\beta\hbar\omega}}{(1 - e^{-\beta\hbar\omega})^2}
 \end{aligned}$$

(b) We start with:

$$\begin{aligned} Z_{AB} &= \sum_{i,j} e^{-\beta E_{ij}} = \sum_{i,j} e^{-\beta E_i^A - \beta E_j^B} = \sum_i \sum_j e^{-\beta E_i^A} e^{-\beta E_j^B} \\ &= \sum_i e^{-\beta E_i^A} \sum_j e^{-\beta E_j^B} = \sum_i e^{-\beta E_i^A} Z_B \\ &= Z_A Z_B \end{aligned}$$

And therefore free energy:

$$F_{AB} = -kT \ln(Z_{AB}) = -kT \ln(Z_A) - kT \ln(Z_B) = F_A + F_B$$

Next, entropy is given as a derivative¹ of F with respect to T - derivatives are linear operations, hence $S_{AB} = S_A + S_B$.

■

¹with some multiplicative constant